### Force, Work, and Energy in Field-Particle Interactions

Mike Reppert

September 7, 2022

#### The Homogeneous Wave Equation:

$$\left(\frac{1}{c^2}\frac{\partial^2}{\partial t^2} - \nabla^2\right)\boldsymbol{e}(\boldsymbol{x},t) = 0.$$



#### The Homogeneous Wave Equation:





#### 2 The Poynting Vector and Energy Density





#### Recall the Lorentz Force Law for finite particles:

$$F_{EM} = q \left( \boldsymbol{e}(\boldsymbol{r}, t) + \boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t) \right)$$

#### Recall the Lorentz Force Law for finite particles:

$$F_{EM} = q \left( \boldsymbol{e}(\boldsymbol{r}, t) + \boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t) \right)$$

The **work** performed by the EM field is the integral of the force over distance:

$$W_{EM} = \int_{\boldsymbol{r}(t_1)}^{\boldsymbol{r}(t_2)} d\boldsymbol{r} \cdot \boldsymbol{F}_{EM}(\boldsymbol{r})$$

#### Recall the Lorentz Force Law for finite particles:

$$F_{EM} = q \left( \boldsymbol{e}(\boldsymbol{r}, t) + \boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t) \right)$$

The **work** performed by the EM field is the integral of the force over distance:

$$egin{aligned} W_{EM} &= \int_{oldsymbol{r}(t_1)}^{oldsymbol{r}(t_2)} doldsymbol{r} \cdot oldsymbol{F}_{EM}(oldsymbol{r}) \ &= \int_{t_1}^{t_2} dt \; oldsymbol{v}(t) \cdot oldsymbol{F}_{EM}(oldsymbol{r}(t)) \end{aligned}$$

# **Notice**: The magnetic field *never* does work on a charged particle!

 $\boldsymbol{v} \cdot (\boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t)) = 0.$ 

**Notice**: The magnetic field *never* does work on a charged particle!

 $\boldsymbol{v} \cdot (\boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t)) = 0.$ 

$$W_{\mathsf{el}} = \sum_{n} q_n \int_{t_1}^{t_2} dt \boldsymbol{v}_n(t) \cdot \boldsymbol{e}(\boldsymbol{r}_n, t)$$

**Notice**: The magnetic field *never* does work on a charged particle!

 $\boldsymbol{v} \cdot (\boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t)) = 0.$ 

$$W_{el} = \sum_{n} q_n \int_{t_1}^{t_2} dt \boldsymbol{v}_n(t) \cdot \boldsymbol{e}(\boldsymbol{r}_n, t)$$
$$= \int_{t_1}^{t_2} dt \int_{V} d\boldsymbol{x} \left( \sum_{n} q_n \boldsymbol{v}_n(t) \delta\left(\boldsymbol{x} - \boldsymbol{r}_n\right) \right) \cdot \boldsymbol{e}(\boldsymbol{x}, t)$$

**Notice**: The magnetic field *never* does work on a charged particle!

 $\boldsymbol{v} \cdot (\boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t)) = 0.$ 

$$\begin{split} W_{\mathsf{el}} &= \sum_{n} q_n \int_{t_1}^{t_2} dt \boldsymbol{v}_n(t) \cdot \boldsymbol{e}(\boldsymbol{r}_n, t) \\ &= \int_{t_1}^{t_2} dt \int_{V} d\boldsymbol{x} \left( \sum_{n} q_n \boldsymbol{v}_n(t) \delta\left(\boldsymbol{x} - \boldsymbol{r}_n\right) \right) \cdot \boldsymbol{e}(\boldsymbol{x}, t) \\ &\approx \int_{t_1}^{t_2} dt \int_{V} d\boldsymbol{x} \boldsymbol{j}(\boldsymbol{x}, t) \cdot \boldsymbol{e}(\boldsymbol{x}, t) \end{split}$$

**Notice**: The magnetic field *never* does work on a charged particle!

 $\boldsymbol{v} \cdot (\boldsymbol{v} \times \boldsymbol{b}(\boldsymbol{r}, t)) = 0.$ 

$$\begin{split} W_{\mathsf{el}} &= \sum_{n} q_n \int_{t_1}^{t_2} dt \boldsymbol{v}_n(t) \cdot \boldsymbol{e}(\boldsymbol{r}_n, t) \\ &= \int_{t_1}^{t_2} dt \int_{V} d\boldsymbol{x} \left( \sum_{n} q_n \boldsymbol{v}_n(t) \delta\left(\boldsymbol{x} - \boldsymbol{r}_n\right) \right) \cdot \boldsymbol{e}(\boldsymbol{x}, t) \\ \end{split}$$
Here Be Dragons!  $\Rightarrow \approx \int_{t_1}^{t_2} dt \int_{V} d\boldsymbol{x} \boldsymbol{j}(\boldsymbol{x}, t) \cdot \boldsymbol{e}(\boldsymbol{x}, t)$ 

The electric field does work on charged particles – the integral of the Lorentz force over the particle displacement.

The magnetic field does no work on charged particles.

For *finite particles*, the EM work can be written as an integral over the current density  $\boldsymbol{j}(\boldsymbol{x},t)$ .

$$W_{\rm el} \approx \int_{t_1}^{t_2} dt \int_V d\boldsymbol{x} \boldsymbol{j}(\boldsymbol{x},t) \cdot \boldsymbol{e}(\boldsymbol{x},t)$$

#### The Poynting Vector and Energy Density

Solving the Maxwell-Faraday equation

$$\nabla \times \boldsymbol{B} - \frac{1}{c} \frac{\partial \boldsymbol{E}}{\partial t} = \frac{4\pi}{c} \boldsymbol{j}(\boldsymbol{x}, t),$$

for  $\boldsymbol{j}(\boldsymbol{x},t)$ 

#### The Poynting Vector and Energy Density

Solving the Maxwell-Faraday equation

$$\nabla \times \boldsymbol{B} - \frac{1}{c} \frac{\partial \boldsymbol{E}}{\partial t} = \frac{4\pi}{c} \boldsymbol{j}(\boldsymbol{x}, t),$$

for  $oldsymbol{j}(oldsymbol{x},t)$ , we get (after some cross-product magic...)

$$W_{\mathsf{el}} = -\int_{t_1}^{t_2} dt \int_V d\boldsymbol{x} \left( \nabla \cdot \boldsymbol{S} + \frac{\partial u}{\partial t} \right)$$

with the Poynting vector

$$oldsymbol{S}(oldsymbol{x},t)\equivrac{c}{4\pi}oldsymbol{e} imesoldsymbol{b}$$

and the *electromagnetic energy density* 

$$u(\mathbf{x},t) = \frac{1}{8\pi} \left( \|\mathbf{e}\|^2 + \|\mathbf{b}\|^2 \right).$$

#### The Poynting Vector and Energy Density

The **energy density** represents the "amount" of electromagnetic energy in a given region of space.

If the volume  $\boldsymbol{V}$  contains the whole field:

$$W_{\mathsf{el}} = -\int_{t_1}^{t_2} dt \int_V d\boldsymbol{x} \, \frac{\partial u}{\partial t} = -\left[\int_V d\boldsymbol{x} \, u(x,t_2) - \int_V d\boldsymbol{x} \, u(x,t_1)\right].$$



#### The Poynting Vector and Energy Density

The **Poynting vector** S represents the *magnitude and direction* of energy flow.

In vacuum,  $\boldsymbol{S}$  and u satisfy a *continuity equation*:

$$\nabla \cdot \boldsymbol{S} + \frac{\partial u}{\partial t} = 0.$$



//www.tf.uni-kiel.de/matwis/amat/admat\_en/kap\_5/backbone/r5\_1\_4.html

Mike Reppert

https:

September 7, 2022 11 / 18

#### Take-Home Points

*Electromagnetic work* can be written as a time and space integral over two quantities:

- The energy density u(x,t) characterizes the "amount" of EM energy in a given region of space
- The Poynting vector  $S(\pmb{x},t)$  characterizes the magnitude and direction of EM energy flow
- The two are related by the continuity equation

$$\nabla \cdot \boldsymbol{S} + \frac{\partial u}{\partial t} = 0.$$



Image: A matrix

æ

Q: How do we measure optical fields experimentally?

## Q: How do we measure optical fields experimentally?

NB: optical fields oscillate too quickly for electronic circuits to follow!

Q: How do we measure optical fields experimentally? NB: optical fields oscillate too quickly for electronic circuits to follow!

A: We measure the *energy absorbed* by a detector.



#### An example: the bolometer

https://cuore.lngs.infn.it/en/about/detectors

Mike Reppert

Energy Metrics for EM fields

**Pulse Energy:** 

$$U_{\mathsf{pulse}} = \int d\boldsymbol{x} \, u(\boldsymbol{x},t).$$



э

Energy Metrics for EM fields

**Pulse Energy:** 

$$U_{\mathsf{pulse}} = \int d\boldsymbol{x} \, u(\boldsymbol{x},t).$$



Energy Metrics for EM fields

Irradiance  $\propto \boldsymbol{S}(\boldsymbol{x},t)\cdot\hat{\boldsymbol{n}}$ :

$$\begin{aligned} & \operatorname{Ir}_{\mathsf{det}} = \frac{c\left(\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}\right)}{4\pi\tau_{\mathsf{det}}A_{\mathsf{det}}} \int_{t_o}^{t_o + \tau_{\mathsf{det}}} dt \int dA \, \|\boldsymbol{e}(\boldsymbol{x}, t)\|^2 \\ &\approx \frac{c\left(\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}\right)}{8\pi^2\tau_{\mathsf{det}}A_{\mathsf{det}}} \int dA \int d\omega \, \|\boldsymbol{\breve{e}}(\boldsymbol{x}, \omega)\|^2 \,. \end{aligned}$$

Determined by  $\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}$  and the **intensity**  $I(\boldsymbol{x}, t)$  or  $I(\boldsymbol{x}, \omega)$ .

Energy Metrics for EM fields

Irradiance  $\propto \boldsymbol{S}(\boldsymbol{x},t)\cdot\hat{\boldsymbol{n}}$ :

$$\begin{split} & \mathsf{Ir}_{\mathsf{det}} = \frac{c\left(\hat{\mathbf{s}}\cdot\hat{\mathbf{n}}\right)}{4\pi\tau_{\mathsf{det}}A_{\mathsf{det}}} \int_{t_o}^{t_o+\tau_{\mathsf{det}}} dt \int dA \, \|\boldsymbol{e}(\boldsymbol{x},t)\|^2 \\ &\approx \frac{c\left(\hat{\mathbf{s}}\cdot\hat{\mathbf{n}}\right)}{8\pi^2\tau_{\mathsf{det}}A_{\mathsf{det}}} \int dA \int d\omega \, \|\boldsymbol{\breve{e}}(\boldsymbol{x},\omega)\|^2 \, . \end{split}$$

Determined by  $\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}$  and the **intensity**  $I(\boldsymbol{x}, t)$  or  $I(\boldsymbol{x}, \omega)$ .



Energy Metrics for EM fields

Irradiance  $\propto \boldsymbol{S}(\boldsymbol{x},t)\cdot\hat{\boldsymbol{n}}$ :

$$\begin{split} & \mathsf{Ir}_{\mathsf{det}} = \frac{c\left(\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}\right)}{4\pi\tau_{\mathsf{det}}A_{\mathsf{det}}} \int_{t_o}^{t_o + \tau_{\mathsf{det}}} dt \int dA \, \|\boldsymbol{e}(\boldsymbol{x}, t)\|^2 \\ &\approx \frac{c\left(\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}\right)}{8\pi^2\tau_{\mathsf{det}}A_{\mathsf{det}}} \int dA \int d\omega \, \|\boldsymbol{\breve{e}}(\boldsymbol{x}, \omega)\|^2 \, . \end{split}$$

Determined by  $\hat{\mathbf{s}} \cdot \hat{\mathbf{n}}$  and the **intensity**  $I(\boldsymbol{x}, t)$  or  $I(\boldsymbol{x}, \omega)$ .





#### "Intensity"

#### Lots of different things get called "intensity"!

				<u> </u>	<b>č</b>
Radiant intensity	l <sub>e,Ω</sub> <sup>[nb 5]</sup>	watt per steradian	W/sr	M-L <sup>2</sup> -T <sup>-3</sup>	Radiant flux emitted, reflected, transmitted or received, per unit solid angle. This is intensity 27/38 ^ V X
Spectral intensity	$I_{e,\Omega,v}^{[nb 3]}$ or $I_{e,\Omega,\lambda}^{[nb 4]}$	watt per steradian per hertz or watt per steradian per metre	W-sr <sup>-1</sup> ·Hz <sup>-1</sup> or W-sr <sup>-1</sup> ·m <sup>-1</sup>	M·L <sup>2</sup> ·T <sup>-2</sup> or M·L·T <sup>-3</sup>	Radiant <b>intensity</b> per unit frequency or wavelength. The latter is commonly measured in W-sr <sup>-1</sup> -mm <sup>-1</sup> . This is a directional quantity.
Radiance	$L_{e,\Omega}^{[nb.5]}$	watt per steradian per square metre	W-sr <sup>-1</sup> -m <sup>-2</sup>	м-т-3	Radiant flux emitted, reflected, transmitted or received by a surface, per unit solid angle per unit projected area. This is a directional quantity. This is sometimes also confusingly called 'intensity'.
Spectral radiance	$L_{e,\Omega,v}$ [nb 3] Or $L_{e,\Omega,\lambda}$ [nb 4]	watt per steradian per square metre per hertz or watt per steradian per square metre, per metre	W-sr <sup>-1</sup> -m <sup>-2</sup> -Hz <sup>-1</sup> or W-sr <sup>-1</sup> -m <sup>-3</sup>	M·T <sup>-2</sup> or M·L <sup>-1</sup> ·T <sup>-3</sup>	Radiance of a surface per unit frequency or wavelength. The latter is commonly measured in W sr <sup>-1</sup> m <sup>-2</sup> nm <sup>-1</sup> . This is a directional quantity. This is sometimes also confusingly called "spectral <b>intensity</b> ".
Irradiance Flux density	Ee <sup>(nb 2)</sup>	watt per square metre	W/m <sup>2</sup>	M-T-3	Radiant flux received by a surface per unit area. This is sometimes also confusingly called "intensity".
Spectral irradiance Spectral flux density	$E_{e,v}^{[rb 3]}$ or $E_{e,\lambda}^{[rb 4]}$	watt per square metre per hertz or watt per square metre, per metre	W·m <sup>-2</sup> ·Hz <sup>-1</sup> or W/m <sup>3</sup>	M·T <sup>-2</sup> or M·L <sup>-1</sup> .T <sup>-3</sup>	Inadiance of a surface per unit frequency or wavelength. This is sometimes also confusingly called "spectral <b>Intensity</b> ". Non-SI units of spectral flux density include junsky (1 Jy = $10^{-20}$ W m <sup>-2</sup> Hz <sup>-1</sup> and Solar flux unit (1 Su = $10^{-20}$ W m <sup>-2</sup> Hz <sup>-1</sup> = $10^{4}$ Jy).
Radiosity	Je <sup>[nb 2]</sup>	watt per square metre	W/m <sup>2</sup>	M-T-3	Radiant flux leaving (emitted, reflected and transmitted by) a surface per unit area. This is sometimes also confusingly called "intensity".
Spectral radiosity	$J_{e,v}^{[nb 3]}$ or $J_{e,\lambda}^{[nb 4]}$	watt per square metre per hertz or watt per square metre, per metre	W·m <sup>-2</sup> ·Hz <sup>-1</sup> or W/m <sup>3</sup>	M·T <sup>-2</sup> or M·L <sup>-1</sup> ·T <sup>-3</sup>	Radically of a surface per unit frequency or wavelength. The latter is commonly measured in W m <sup>-2</sup> nm <sup>-1</sup> . This is sometimes also confusingly called "spectral <b>intensity</b> ".
Radiant exitance	Me <sup>(nb 2)</sup>	watt per square metre	W/m <sup>2</sup>	м-т-3	Radiant flux emitted by a surface per unit area. This is the emitted component of radiosity. 'Radiant emittance' is an old term for this quantity. This is sometimes also confusingly called 'intensity'.
Spectral exitance	M <sub>e,V</sub> [nb 3] or M <sub>e,X</sub> [nb 4]	watt per square metre per hertz or watt per square metre, per metre	W·m <sup>-2</sup> ·Hz <sup>-1</sup> or W/m <sup>3</sup>	M-T <sup>-2</sup> or M-L <sup>-1</sup> .T <sup>-3</sup>	Radiart extrance of a surface per unit frequency or wavelength. The latter is commonly measured in W·m <sup>-2</sup> nm <sup>-1</sup> , "Spectral emittance" is an old term for this quartity. This is sometimes also confusingly called "spectral <b>intensity</b> ".
Radiant exposure	He	joule per square metre	J/m <sup>2</sup>	M-T-2	Radiant energy received by a surface per unit area, or equivalently irradiance of a surface integrated over time of irradiation. This is sometimes also called "radiant fluence".

#### https://en.wikipedia.org/wiki/Intensity\_(physics)

Mike Reppert

#### Take-Home Points

Optical fields are characterized experimentally by the energy they carry.

Detectors monitor EM fields by measuring the *energy absorbed* by optically dense materials.

The *pulse energy* refers to the total EM energy carried by an ultrafast pulse, i.e.  $\int dx \, u(x,t)$ 

Irradiance refers to the rate at which a beam transmits energy in a given direction, i.e.  $S\cdot\hat{n}$ 

Informally, we use the term "intensity" for either  $I(\boldsymbol{x},t) = \|\boldsymbol{e}(\boldsymbol{x},t)\|^2$  or  $I(\boldsymbol{x},\omega) = \|\breve{\boldsymbol{e}}(\boldsymbol{x},\omega)\|^2$ .